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Testing of the Prototype Facets for the Stretched-Membrane Faceted Dish

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TESTING OF THE PROTOTYPE FACETS FOR THE STRETCHED-MEMBRANE FACETED DISH

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Abstract

The Faceted Stretched-Membrane Dish Program is part of a DOE-sponsored effort to develop a commercial 25 kWe dish/Stirling system employing a twelve-facet dish concentrator. The facets will utilize the stretched-membrane technology originated in the heliostat development program. Each facet is constructed with a thin metal membrane stretched over both sides of a steel ring. When a small vacuum is induced between the membranes they assume a parabolic contour capable of concentrating sunlight at a predetermined focal length. A reflective polymer film is attached to the face of the facet to enhance the optical performance.

During Phase II of the Faceted Stretched-Membrane Dish Program, Science Applications International Corp. and Solar Kinetics, Inc., constructed prototype 3.5-meter facets utilizing different design approaches to demonstrate their manufacturability and optical performance. Sandia engaged in a program to determine the on-sun performance of the facets (for f/D s of 2.7 to 3.0). A uniformly distributed slope error was used as the basis for comparison. Flux arrays based on slope error from a computer model were compared to a measured flux array for each facet. The slope error for the facet was determined by the value that would produce a modeled array with the minimum mean square difference to the measured array. The facet produced by SAIC demonstrated uniform slope errors of 2.2 to 3.0 milliradians with peak flux intensities of 334 to 416 kW/m². The SKI facet had slope errors of 1.6 to 1.9 milliradians with peak flux intensities of 543 to 1186 kW/m².

1. Introduction

1.1 Test Facility

The National Solar Thermal Test Facility (NSTTF) [1,2] is operated by Sandia National Laboratories for the U. S. Department of Energy's Solar Thermal Program. The NSTTF, located on Kirtland Air Force Base in Albuquerque, NM, is capable of supporting a wide range of solar experiments. Data acquisition, control and

diagnostic systems are available for the use of the experimenter. In addition, specialized instrumentation such as the Beam Characterization System (BCS), used to measure the solar intensity profiles produced by all types of concentrators, has been developed at the NSTTF. Sandia has also developed several computer codes, HELIOS [3] and CIRCE[4] and CIRCE2 [5], to model solar concentrator and receiver optical performance. Incorporated in the analysis of the results is a methodology for analytically comparing the real image measurements with the theoretical predictions, and it enhances our ability to predict receiver performance.

1.2 Faceted Stretched-Membrane Dish Program

The goal of the Faceted Stretched-Membrane Dish Program is to develop a 25 kW_e modular dish/Stirling power-production system. The stretched-membrane facets are similar in concept to the stretched-membrane heliostat. In general, the facets are made by attaching tensioned metal membranes to a steel support ring. One of the membranes has a reflective surface on the exposed face. When a small vacuum is drawn on the inside of the structure, the membranes assume a concave shape, and the reflective surface acts as a focusing mirror. One design discussed in this report applies a different approach, employing a plastically deformed front membrane. Twelve of the facets will be mounted on a space frame to form a parabolic concentrator, as shown in Figure 1. The design goal for the facet uniform slope error is 2.5-mr (1 standard deviation)[6].

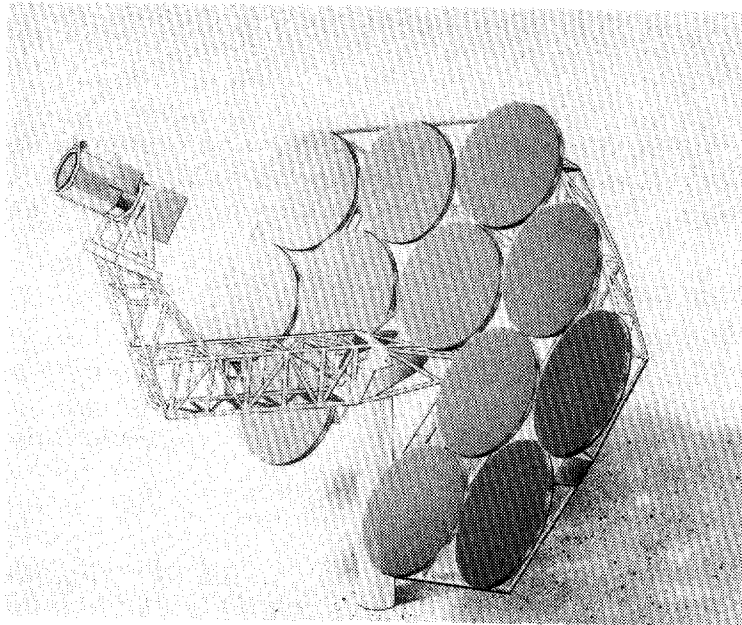


Figure 1. Stretched-Membrane Faceted Dish Conceptual Drawing.

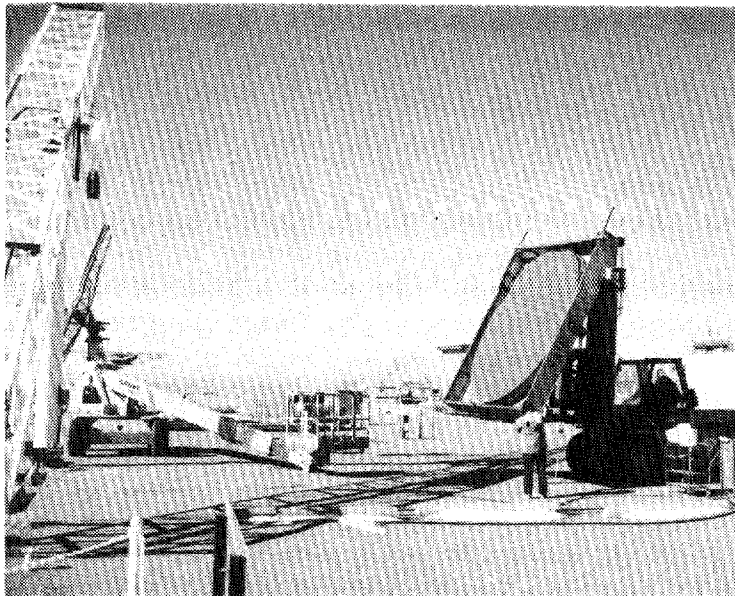


Figure 2 On-Sun Testing of a Stretched-Membrane Facet

As part of the program, two contractors, Science Applications International Corp. (SAIC)[7] and Solar Kinetics, Inc., (SKI)[8] fabricated prototype 3.5-m. diameter facets that were tested on-sun at the NSTTF (Figure 2.). This report summarizes the testing and analysis done in connection with this phase of the program.

2. Test Facets

2.1 The SAIC Facet

The membrane for the facet supplied by SAIC was fabricated from 0.076-mm thick, Type 201, half-hard stainless steel. The membranes are pre-tensioned and then welded to the facet ring. The front membrane has a laminated reflective surface of silvered EPC 305, a 3M company product. Aluminized EPC 244 tape was used to seal the exposed seams of the laminate. Focusing of the SAIC facet is achieved by providing a vacuum in the space between the membranes. The membranes deflect inward until the front one contacts

the internal focus control valve and shuts off the vacuum. The deflections of the front membrane from the ring plane necessary to meet the required focal lengths were supplied by SAIC (Table 1). The position of the focus control valve was adjusted until the proper displacement was measured at the center of the front membrane.

2.2 The SKI Facet

SKI uses a different approach in fabricating its prototype facet. The front membrane is plastically deformed during fabrication to a contour that approximates a parabola. The front membrane of the prototype is 0.08-mm thick 304 stainless steel. The reflective film laminated on the front membrane is EPC 305. The rear membrane is a 304 stainless steel membrane 0.10-mm thick. At the center of the facet, a tether is installed between the membranes to stabilize the thin front surface in winds. The membranes are held to the ring assembly with spring clips.

To focus the facet, a pressure difference is maintained between the front face of the facet and the interior. A pressure controller is used to cycle a vacuum pump to sustain the prescribed differential. Set

points for the controller (Table 2) were provided by SKI.

3. Test Equipment and Configuration

3.1. Beam Characterization System

The test instrumentation used to measure the on-sun flux density distributions produced by concentrators is Sandia's Beam Characterization System (BCS). A schematic diagram of the system is shown in Figure 3. The BCS comprises a lambertian target plane with an internal flux gauge, a video camera with neutral density filters, and a computer system with a frame grabber that digitizes the video image, displays it on a monitor and stores it for later evaluation. The digitized image is processed with a software package called Beamcode® (enhanced for Sandia by the developer Big Sky Software) to provide flux contours, total beam power, and the flux-density distributions.

The solar image from the concentrator is reflected onto a water-cooled target with a plasma-sprayed aluminium oxide coating. Previous laboratory measurements of these surfaces have shown them to be nearly lambertian [9]. Flux gauges are located at two points in the target surface, providing a direct measurement of the flux density at their location within the reflected flux-density profile so that the measured gray-scale level can be scaled and equated with a solar intensity. The flux gauge configuration is a water-cooled circular foil heat flux gauge.

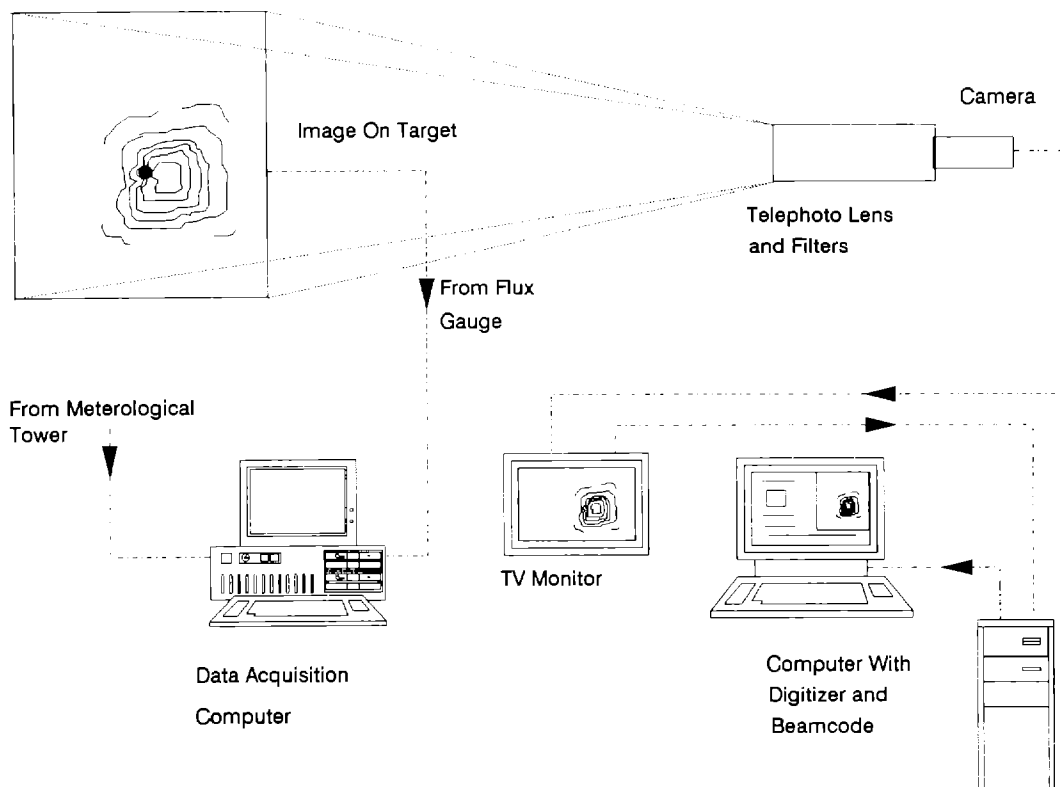


Figure 3. Beam Characterization System Schematic

A video camera is used to view the reflected image at the receiver plane. The camera has a visual spectral response, a wide dynamic range, high resolution, zero geometric distortion, and no lag or image retention. Standard "C" mount lenses with neutral density filters are used on the camera to adjust the intensity levels viewed by the camera to avoid saturation. The video frame grabber and digitizer provides a spatial resolution up to 240 by 240 pixels with 256 grey-scale intensity levels. Images can be captured at rates up to 60 per second.

3.2 Test Configuration

Both facets were delivered to Sandia mounted in shipping containers that also served as the facet support structure. A mounting frame for the containers was built at the NSTTF that could hold the facet at an angle on a fork lift. The fork lift used for positioning the facets has the capability for six directions of motion (up-down, forward-backward, left-right, tilt, pan, and rotation). A tarp sized to cover most of the face of the facets was fixed to the frame so it could be raised and lowered to shield the facets from the sun when focusing of the sun was not desired.

A water-cooled target with flux gauges was mounted at a fixed position in an open area of the test site. A measurement grid was surveyed and marked on the asphalt north of the target. The facets were positioned on the grid with the fork lift so that the concentrated beam struck the target. The BCS camera was positioned adjacent to the facet to record the images. Figure 4 is a schematic of the test configuration.

3.3 Uncertainties In The Measurements

The uncertainties in the measurements made during on-sun testing are listed in Table 3. The most important reduced measurement made in these tests is the scale factor that was developed to establish the flux levels associated with the measured flux-density distributions. This measurement is the first uncertainty in the chain of measurements that leads to a flux uncertainty. Also part of this chain is the uniformity of the neutral density filters, the correction for the angle between the camera view and the target normal, and the uncertainty in the size and intensity of a pixel in the final distribution. Our estimate is that the measured flux-densities are

within $\pm 8\%$ of the reported peak-flux value.

4. Test Procedure

For each test, the focus control was set according to the instructions supplied by the facet fabricator. The covered facet was positioned with the fork lift on the grid so that the reflected beam was striking the target. With the tarp lowered, enough of the sides of the facet were uncovered to allow initial positioning of the facet.

Measurements from the bottom edge of the facet mounting frame to the grid provided input to calculations to determine the vertex to target distance and azimuth angle. The facet position was adjusted until the calculated distance was close to the desired focal length, and the facet was uncovered to allow capture of a series of BCS images. This process was repeated for both facets at each focal length. The second and third columns of Table 4 show the desired (design) and the actual (test) focal lengths. Details of the calculations are in Appendix 1.

The object of these tests was to measure the optical performance of the facets at fixed focal lengths corresponding to the positions of the facets on the dish. The test setup did not have any automatic sun-tracking capability. The facets were manually positioned for each test, and the procedure was not intended to bracket the focal point or determine the optimal focal length. Since the completion of these tests, a test apparatus equipped with sun tracking and a movable target has been developed and will be used to test the next generation of facets.

5. Test Results

The images captured with the BCS were analyzed to determine peak flux and image shape. Peak flux results are listed in the fourth column of Table 4. Flux values reported by Beamcode[®] are relative to the zero background of the image. As part of the test procedure for the BCS, a nonilluminated image of the target is made prior to the actual test sequence. This image is subtracted from each test image by the software as a means of reducing "background noise." As a result, each pixel in the digitized test image has an integer value of 0 to 255 relative to the 0 pixel level. Actual flux values are determined using the flux gauge measurements and

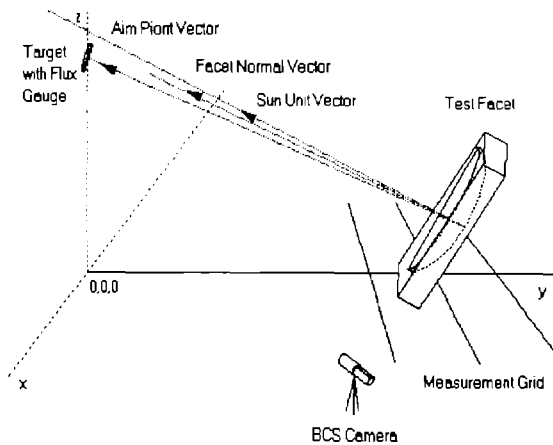


Figure 4. Stretched-Membrane Facet Test CIRCE2 Configuration

Table 3
Measurement Uncertainties

Measurement Variable	Measured Uncertainty
Time of Day	± 5 Seconds
Ambient Temperature	± 0.28 °C
Wind Speed	± 5 % of reading
Wind Direction	± 10 % of reading
Solar Radiation	± 2 % of reading
Target Flux	± 8 % of maximum value
Target-to-Facet Distance	± 0.03 m
Azimuth and Elevation angles	± 0.5 Degrees
Reflectivity	± 2 % of reading

are normalized to 1 kW/m². Flux gauge locations in the target appear as holes in the image surface. Viewing profiles of the hole with the software provide a means for gauging the relative intensity of the pixels at the edge of the gauge location. By interpolation, a relative intensity for the pixel located at the center of the gauge can be determined. This pixel value is ratioed with the measured flux from the gauge to determine the intensity scale factor (kW/m²/pixel relative intensity) for the entire image array. The peak flux values are the average of the scale factor for each flux gauge times the peak pixel relative intensity of the image. The power (fifth column Table 4) is determined by multiplying the sum of the image array by the intensity scale factor and the amount of the target area in one pixel.

Table 4
Stretched-Membrane Facet Test Results

Test Number	Design Focal Length m	Test Focal Length m	Peak Flux kW/m ²	Power kW
SAIC				
H0121144	10.5	10.7	416	8.50
H0121213	10.5	10.5	338	8.79
H0121224	10.5	10.6	389	9.33
H0151209	10.0	10	342	8.68
H0151223	10.0	10.1	357	8.71
H0151232	10.0	10.0	339	8.59
H0231122	9.5	9.6	355	8.92
H0231136	9.5	9.5	334	9.20
H0231148	9.5	9.6	346	8.92
H0231201	9.5	9.6	347	9.12
H0231210	9.5	9.7	365	9.16
H0231218	9.5	9.7	368	9.04
SKI				
H0251155	10.5	10.4	784	9.17
H0251205	10.5	10.6	798	9.16
H0251212	10.5	10.6	843	9.52
H0251220	10.5	10.6	638	9.21
H0251228	10.5	10.7	797	9.57
H0281129	10.0	10.1	1163	8.53
H0301145	10.0	10.3	1090	8.54
H0301154	10.0	10.0	1186	9.08
H0301208	9.5	9.6	543	7.95
H0301218	9.5	9.5	755	7.99
H0311031	9.5	9.8	645	6.19
H0311039	9.5	10.0	921	8.82
H0311121	9.5	10.0	1182	8.86

During testing, some problems were encountered with the focusing system for the SKI facet. Covering and uncovering the facet with the tarp caused large pressure fluctuations that the control system could not accommodate. While the test conditions could have exaggerated the pressure changes, whenever cloud transients occur they may also induce pressure changes which might affect the control system. In addition, while the prescribed vacuum differential for the 9.5-m focal length could be maintained, the actual focal length could not be established. The vertex-to-target distance was varied while using the BCS to determine the beam size. The smallest diameter image corresponded to a vertex-to-center distance of 9.9 m. The measured peak flux (1182 kW/m²) was also consistent with the values measured at 10 m. This would indicate that the change in vacuum did not cause a corresponding change in focal length or facet contour, and the decrease in peak power at the 9.5-m. focal length was a function of the movement away from the 10-m. focal point.

6. Test Analysis

6.1 The CIRCE2 Computer Code

CIRCE2 is a dish/receiver- specific code adapted from HELIOS, a code developed in the late 1970's for central receiver systems (in Greek mythology Circe is the daughter of Helios). The solution technique employed in CIRCE2 is the same as in HELIOS; that is, the concentrator errors are convolved with the sunshape to produce the flux density distribution on an arbitrary target plane.

6.2 CIRCE2 Input

Input to CIRCE2 is a file containing the parameters listed below arranged in a specific sequence. A short program creates the file and prompts the user for the required input.

Sunshape

The sunshape input can be either Gaussian, a uniform disk with any of six limb-darkening options, or a user-specified tabular input. A clear-day, tabular sunshape, measured at the NSTTF, was used for these calculations.

Sun, Dish, and Target Orientation

In CIRCE2, the default position of the sun is directly overhead of an upward facing dish. Inputs to the program allow for specification of the position vectors for the sun, the target, and the concentrator. By transforming the test coordinate system to CIRCE2 coordinates, the program inputs model the actual positions of the sun, target and concentrator. Appendix 1 details the calculations for determining the position inputs to CIRCE2.

Error Parameters

Up to five different reflector errors can be input to the code as either one-dimensional (circular normal) or two-dimensional (elliptic normal) errors. A single, circular-normal slope error was used to model the performance of the facets. The slope error is input to the code to model flux distribution. The slope error is varied until predicted peak flux closely matches the measured peak flux. This match is determined when a 0.1 change in the slope error determines flux values that bracket the measured flux value.

Convolution

The convolution of the errors and the sunshape can be one- or two-dimensional and either numerically or analytically calculated. A two-dimensional numerical convolution is used for the concentrators because of the offsets and the tabular sun input.

Target Shadowing

Target shadowing or blockage of the reflective surface can be input as a percentage of the concentrator projected area, or computed internally by overlaying a projection of the target on the facets. Target shadowing is neglected in these tests since no shadowing occurred.

Reflector Types

CIRCE2 can model either continuous surfaces or faceted concentrators. The reflectance of the optical surface is also an input variable. The facets were modeled as continuous surfaces, and the measured solar reflectivity was input to the models.

Facet Shape

CIRCE can support a number of different reflector shapes. For these calculations, circular was used.

Facet Contour

The facet contours were modeled as parabolic.

6.3 CIRCE2 Slope Error

The CIRCE2 code models the facet as a contour of revolution; that is, the reflector surface is axisymmetric. The primary assumption in the code is that the slope errors are uniformly distributed over the surface of the reflector. In fact, this is rarely the case since fabrication techniques often result in organized departures of the facet contour from design. Nonetheless, the uniformly distributed slope error used in CIRCE2 is useful as a figure-of-merit for comparing the relative performance of the facets.

6.4. Comparing Images

One method for using CIRCE2 and the BCS for analyzing test results has been to vary the uniform slope error in CIRCE2 until the peak flux determined by the code closely matches the value ascertained from the image data. The BCS software allows input of a factor to scale the image to the actual peak flux and includes

some functions to provide graphical and digital information. However, the size of the files is much larger than the 25 by 25 flux intensity array generated by CIRCE2, so visual comparisons of the image contours or profiles for analysis have been the general practice. This is a fairly subjective method, which uses uniform slope error as a figure-of-merit usually reported with two significant digits.

A second more rigorous comparison method has been applied to the analysis. The comparison begins by converting the binary image data file, which represents the entire viewing area of the digitizer, to a 240 by 240 array of real intensities. The actual beam image is only a portion of this array. The BCS software is used to locate the centroid of the image, determine the size of the pixels, and define a rectangular aperture enclosing the image. This information can be used to extract a 25 by 25 flux intensity array which is directly comparable to the CIRCE2 generated array.

The dimensions of the aperture used to extract the actual image array are used as the input for the target dimensions in CIRCE2. The slope error determined by the code that matches the peak power of the image data is used as the starting point for a series of additional CIRCE2 runs. In each run (a minimum of four are done), the slope error is varied by half a milliradian. The series is determined so that the starting slope error is not one of the endpoints of the series.

The flux intensity arrays calculated in the series are put into a spreadsheet along with the extracted measured intensity array. Each array is normalized by its maximum value. For each slope error increment, an array of the differences between the measured and calculated values is created. A statistical figure-of-merit for the goodness of the fit is then calculated by computing the square root of the sum of squares (RSS) of the elements in each difference array.

7. Analysis Results

One set of test results for both facets at each focal length was analyzed with CIRCE2. The facet location measurements were used to calculate facet normal and aim point vectors for input into the program (see Appendix 1). Sun position was determined from the time and date. The coordinate system used is shown in Figure 4.

The procedure used in the analysis began with a slope error determination based on peak power. This provided a starting point for calculating 0.5 mr interval arrays for use in the RSS difference. For each test, a graph of these RSS differences versus slope error was used to determine the calculated slope error with the least RSS difference. Figure 5 is an example of these graphs. The data shown is for the SAIC facet at the 10-m focal length (test #H0151223). The data is contained in Appendix 2 and includes these curves as well as contour plots of the actual image and the CIRCE2 output. A fourth order polynomial curve fit is used to generate the line connecting the data points. The analysis results are summarized in Table 5. Included with the results are uniform slope errors measured at the National Renewable Energy Laboratory (NREL formerly SERI) prior to the Sandia tests using the SHOT [10] system. The SHOT measurements were made under laboratory conditions using a laser ray trace system. Figure 6 shows the RSS difference results compared with

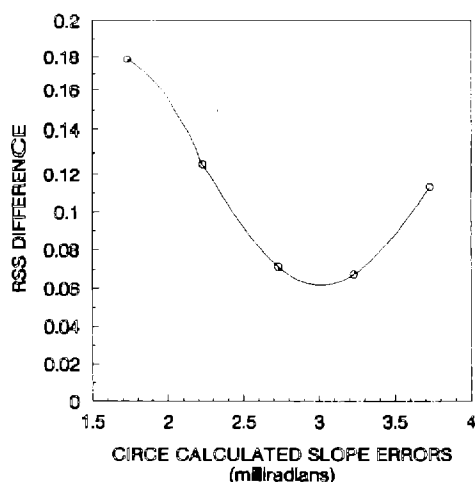


Figure 5 RSS Difference

Table 5.
CIRCE2 And SHOT Results For The Stretched-Membrane Facets

Test #	Focal Length (m)	Peak Flux Matching Slope Error (mr)	RSS Difference Slope Error (mr)	SHOT Slope Error (mr) (Focal Length)
SAIC				
H0121144	10.5	2.2	2.2	2.6(10.45)
H0151223	10.0	2.8	3.0	2.7(9.9)
H0231218	9.5	2.8	3.0	2.8 (9.6)
SKI				
H0251212	10.5	1.5	1.6	1.5 (10.45)
H0301154	10.0	1.3	1.2	1.2 (9.9)
H0301218	9.5	2.0	1.9	1.3 (9.6)

the SHOT measurements. The close agreement between these completely independent measurements reinforces the validity of all the results as they apply to these facets.

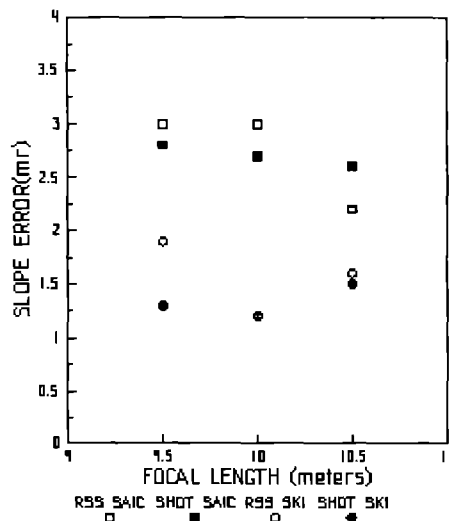


Figure 6 Comparison of RSS Difference SHOT Slope Errors

8. Conclusions

1. Slope errors for the SAIC facet are slightly higher than the design goal of 2.5 mr and are considered acceptable for this stage of the project. The focus control system functions properly and maintains set point.
2. The SKI facet slope errors surpass the design goal for optical performance. However, the focus control system did not perform in a completely predictable manner and requires more development.
3. The on-sun slope error results agree well with independent measurements made with SHOT.

The results of the analysis indicate that, for prototypes, both designs did well and are approaching, if not already capable of meeting, the design goal for optical performance. Both facets are still in the development stage and the lessons learned during these tests will help improve the final designs.

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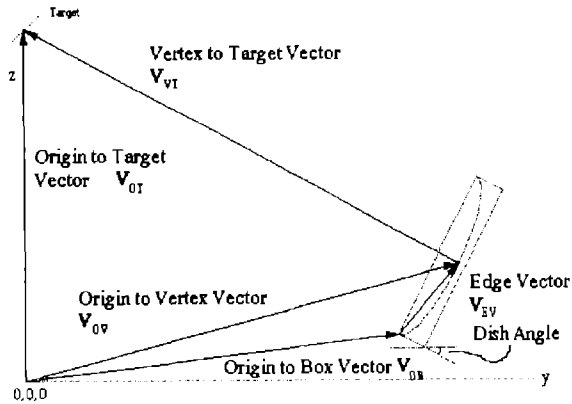
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10. Appendices

Appendix 1
Equations

Determining Vertex to Target Distance



$$|V_{VT}|^2 = |V_{OT}|^2 + |V_{OV}|^2 - 2|V_{OT}|*|V_{OV}|\cos(\theta)$$

where :

$$|V_{OT}|*|V_{OV}|\cos(\theta) = V_{OT} \bullet V_{OV}$$

and

$$V_{OT} = V_{OB} + V_{EV}$$

Grid Layout Centerpoint Dimensions (Distance From Origin)

Centerpoint	X_{Ln}	Y_{Ln}	Z_{Ln}
First Line L1	0	7.94 m	-0.21 m
Second Line L2	0	8.39 m	-0.21 m
Third Line L3	0	8.85 m	-0.21 m

During testing, distances were measured from the box edge(center and corners) to the centerpoint of:

L1 for the 9.5-m focal length (center coordinates X_1, Y_1, Z_1).

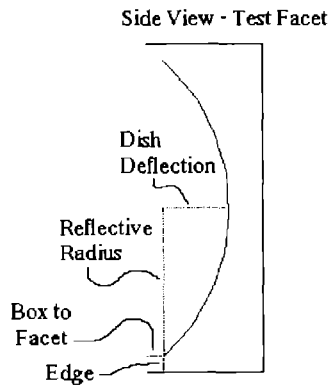
L2 for the 10.0-m focal length (center coordinates X_2, Y_2, Z_3).

L3 for the 10.5-m focal length (center coordinates X_3, Y_3, Z_3).

The Origin to Box Vector V_{OB} is determined by adding the grid dimensions to the measured distances:

$$V_{OB} = (X_n + X_{Ln}) + (Y_n + Y_{Ln}) + (Z_n + Z_{Ln})$$

The Edge Vector V_{EV} is determined from the test facet dimensions and the Dish Angle (measured with a clinometer before each test):



Test Facet Dimensions

Dimension	SAIC meters	SKI meters (vacuum in water)
@ 9.5-m	0.0833	0.0833 (3.95)
@ 10.0-m	0.0800	0.0818 (3.2)
@ 10.5-m	0.0751	0.0781 (2.54)
Reflective Radius	1.861	1.814
Box to Facet	0.109	0.044
Edge	0.134	0.206

$$\mathbf{V}_{EV} = \mathbf{X}_{EV} + \mathbf{Y}_{EV} + \mathbf{Z}_{EV}$$

where :

$$\mathbf{X}_{EV} = 0$$

$$\mathbf{Y}_{EV} = (\text{Box to Facet} + \text{Dish Deflection}) \cos(\text{Dish Angle}) \\ + (\text{Reflective Radius} + \text{Edge}) \sin(\text{Dish Angle})$$

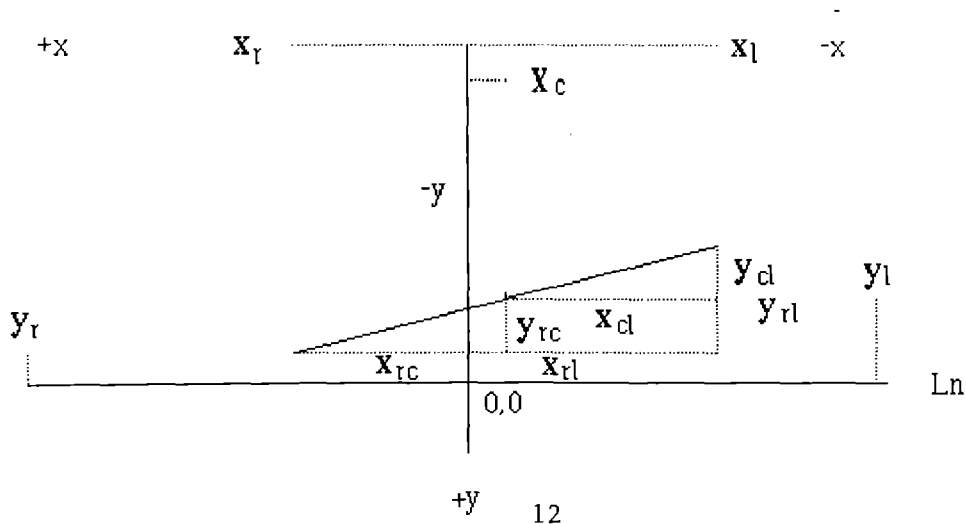
$$\mathbf{Z}_{EV} = (\text{Reflective Radius} + \text{Edge}) \cos(\text{Dish Angle}) \\ - (\text{Box to Facet} + \text{Dish Deflection}) \sin(\text{Dish Angle})$$

\mathbf{V}_{OB} and \mathbf{V}_{EV} are added to obtain \mathbf{V}_{OV} . As shown above the magnitude of \mathbf{V}_{VT} (vertex to target distance) can be determined from \mathbf{V}_{OT} and \mathbf{V}_{OV} . By substituting into the equation, the value for the vertex to target distance becomes:

$$|\mathbf{V}_{EV}| = \sqrt{(X_{OT})^2 + (Y_{OT})^2 + (Z_{OT})^2 + (X_{OV})^2 + (Y_{OV})^2 + (Z_{OV})^2 - 2(X_{OT} * X_{OV} + Y_{OT} * Y_{OV} + Z_{OT} * Z_{OV})}$$

Determining Azimuth Angle

The azimuth angle is determined from the position of the box edge with respect to the ground grid in the x-y plane. The measurements used are all relative to the centerpoint of the grid line.



$$x_{rc} = x_r - x_c \quad y_{rc} = y_r - y_c$$

$$x_{cl} = x_c - x_l \quad y_{cl} = y_c - y_l$$

$$x_{rl} = x_r - x_l \quad y_{rl} = y_r - y_l$$

$$\tan(\theta) = \frac{|y_{rc}|}{|x_{rc}|} = \frac{|y_{cl}|}{|x_{cl}|} = \frac{|y_{rl}|}{|x_{rl}|}$$

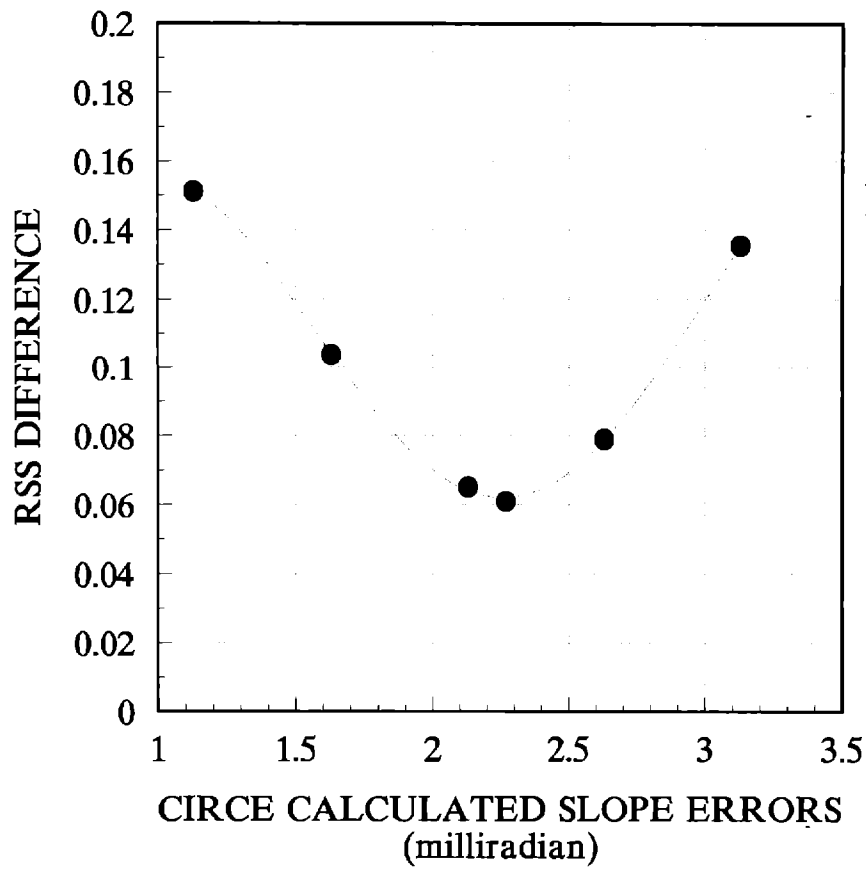
$$\text{Azimuth angle} = 90^\circ + \theta$$

After each test, a measurement of the vertex-to-target distance was made by stretching a steel cable from the target to the center of the covered facet. This measurement was made to verify the above calculations and did not include any azimuth angle measurement. The table below summarizes these measurements. The maximum difference between the values is 0.2 m. The last three tests do not indicate a design focal length. They were

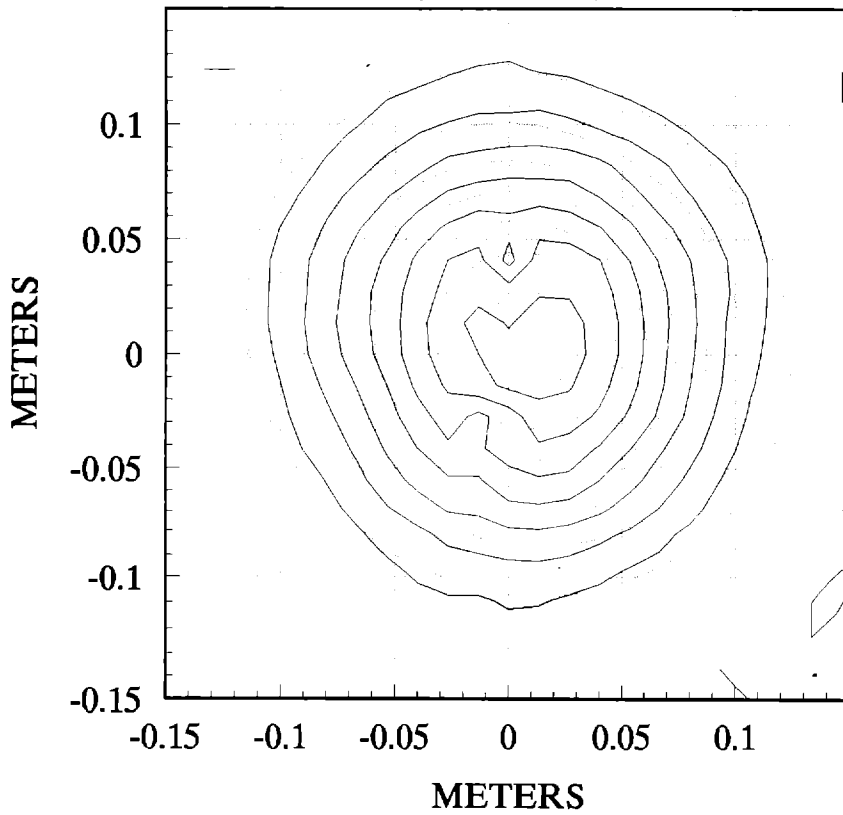
done during the effort to evaluate the 9.5-m focal length discussed in section 5 of the report.

Test Number	Design Focal Length m	Measured Distance m	Calculated Distance m
SAIC			
H0121144	10.5	10.7	10.7
H0121213	10.5	10.5	10.5
H0121224	10.5	10.6	10.6
H0151209	10.0	9.9	10.0
H0151223	10.0	10.0	10.1
H0151232	10.0	10.0	10.0
H0231122	9.5	9.5	9.6
H0231136	9.5	9.4	9.5
H0231148	9.5	9.5	9.5
H0231201	9.5	9.6	9.6
H0231210	9.5	9.6	9.6
H0231218	9.5	9.7	9.7
SKI			
H0251155	10.5	10.3	10.4
H0251205	10.5	10.5	10.6
H0251212	10.5	10.5	10.5
H0251220	10.5	10.6	10.6
H0251228	10.5	10.6	10.7
H0281129	10.0	10.1	10.1
H0301145	10.0	10.3	10.3
H0301154	10.0	10.0	10.0
H0301208	9.5	9.5	9.6
H0301218	9.5	9.5	9.5
H0311031	9.5	na	9.8
H0311039	9.5	na	10.0
H0311049	na	10.4	10.6
H0311109	na	na	9.6
H0311121	na	9.9	10.0

Appendix 2

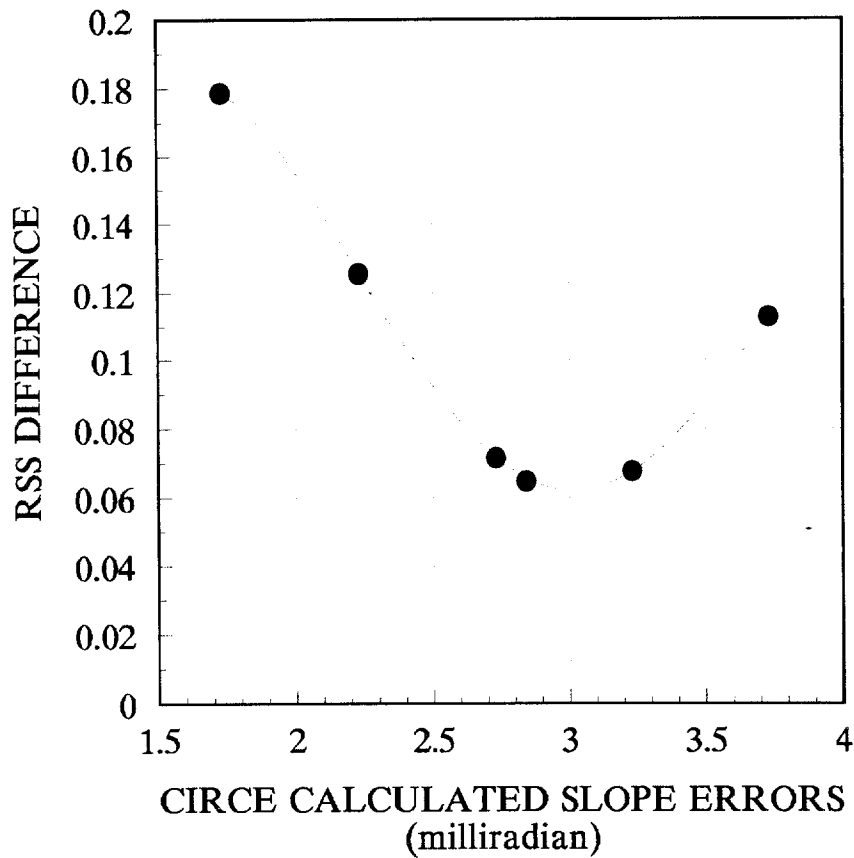


SAIC 10.5 meters
 Peak Power 416 kW/sqm
 Peak Power Method Slope Error 2.3 mr.
 RSS Difference Method Slope Error 2.3 mr.

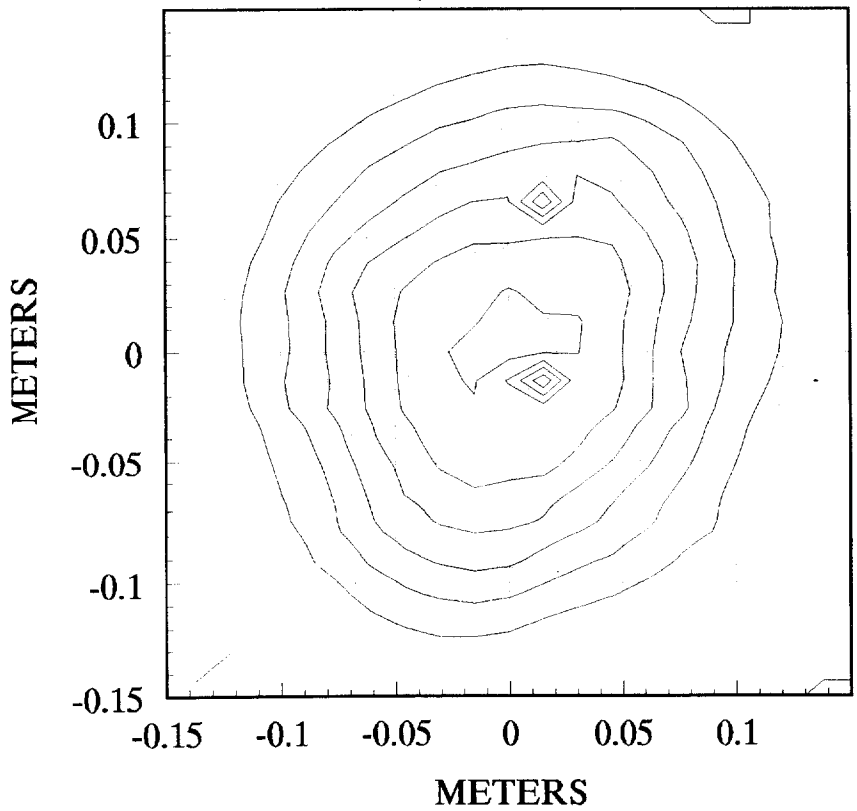


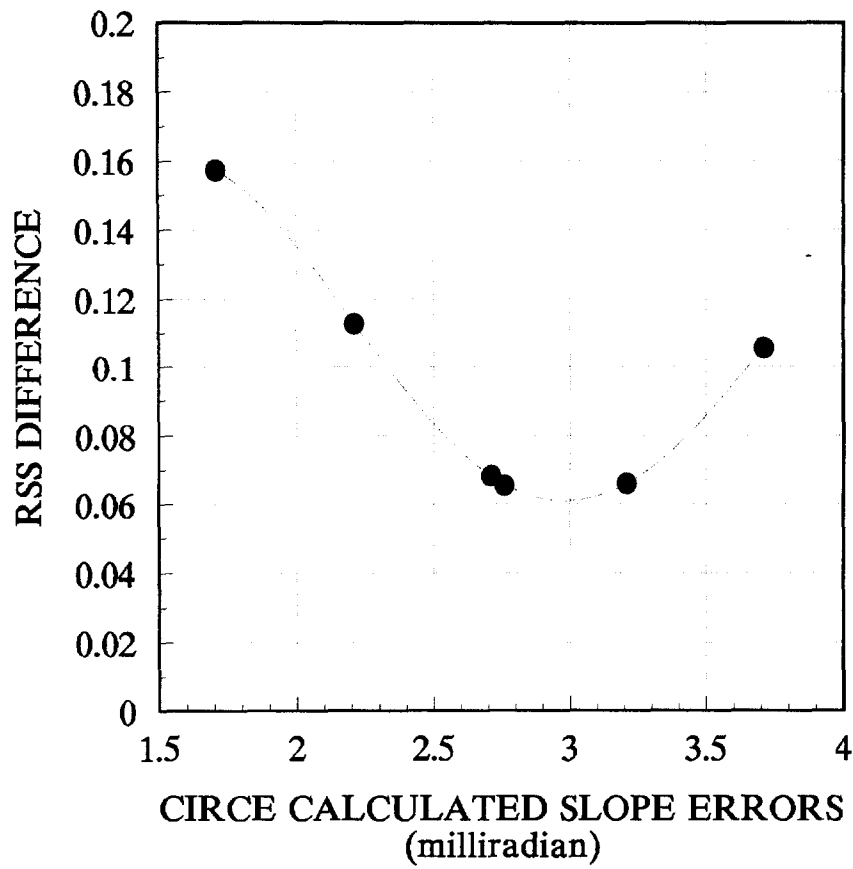
— Actual Image
 — CIRCE2 Prediction

50 Suns/Contour

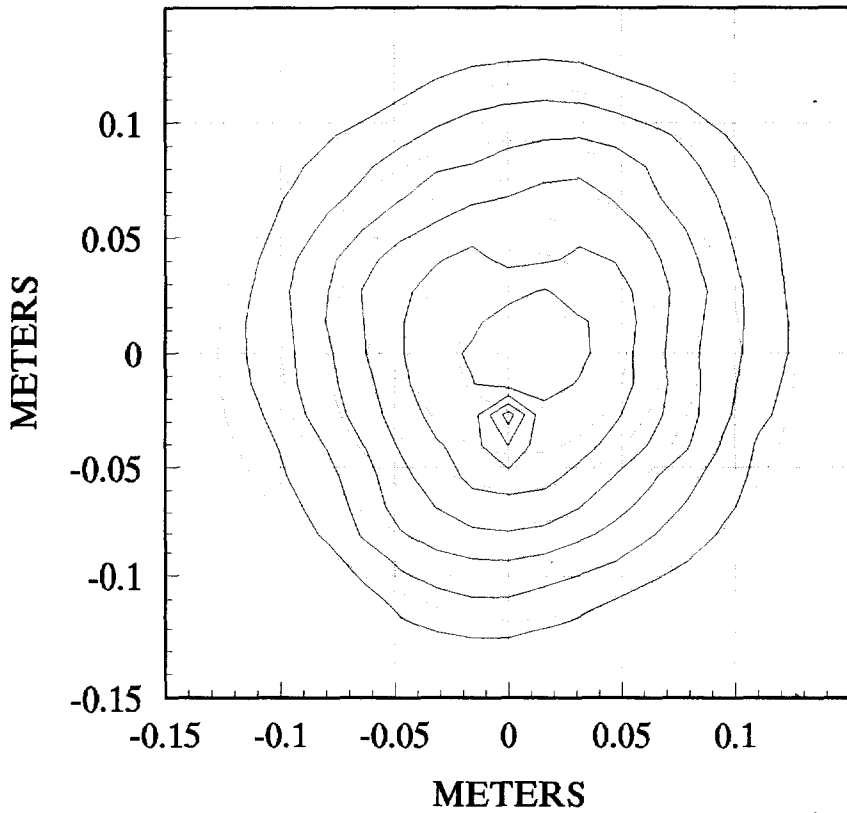


SAIC 10 meters
 Peak Power 357 kW/sqm
 Peak Power Method Slope Error 2.8 mr.
 RSS Difference Method Slope Error 3.0 mr.



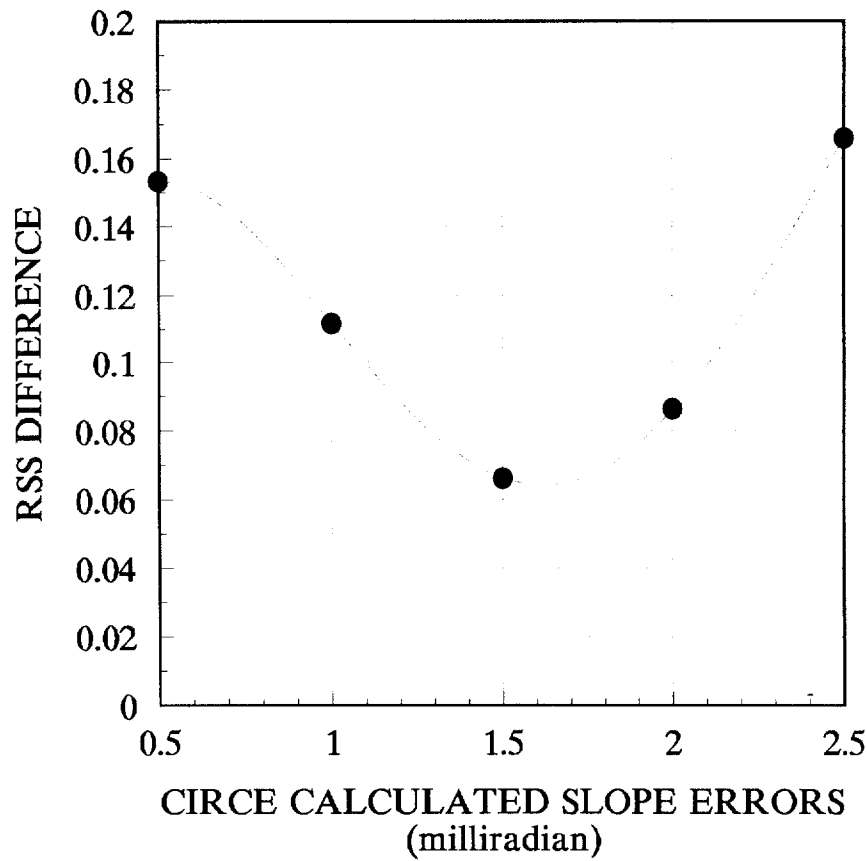


SAIC 9.5 meters
 Peak Power 368 kW/sqm
 Peak Power Method Slope Error 2.8 mr
 RSS Difference Method Slope Error 3.0 mr

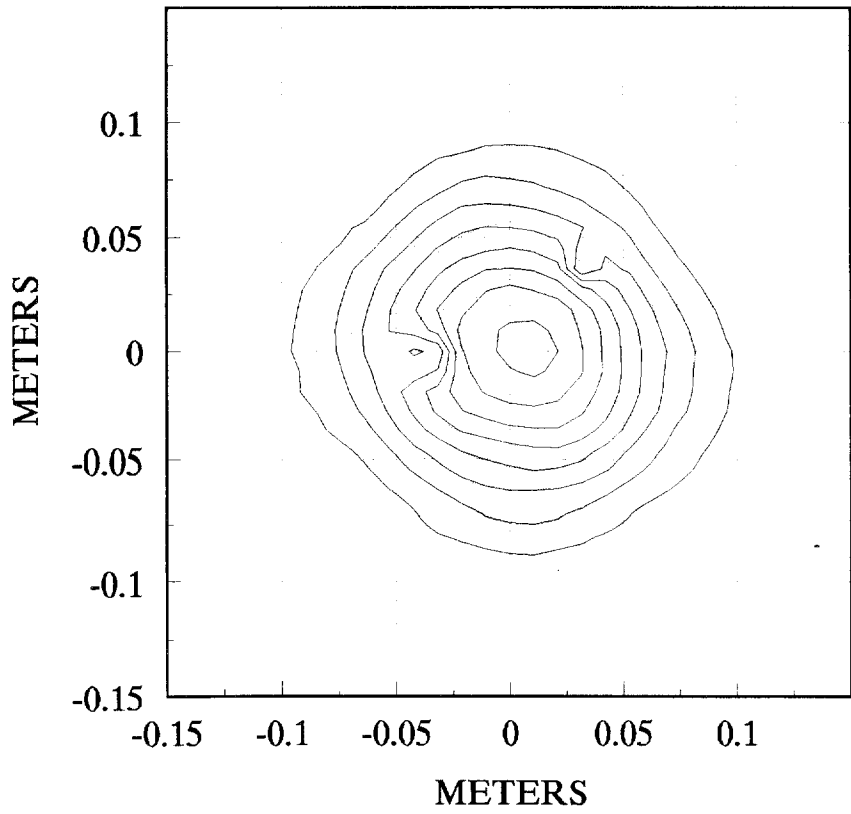


— Actual Image
 — CIRCE2 Prediction

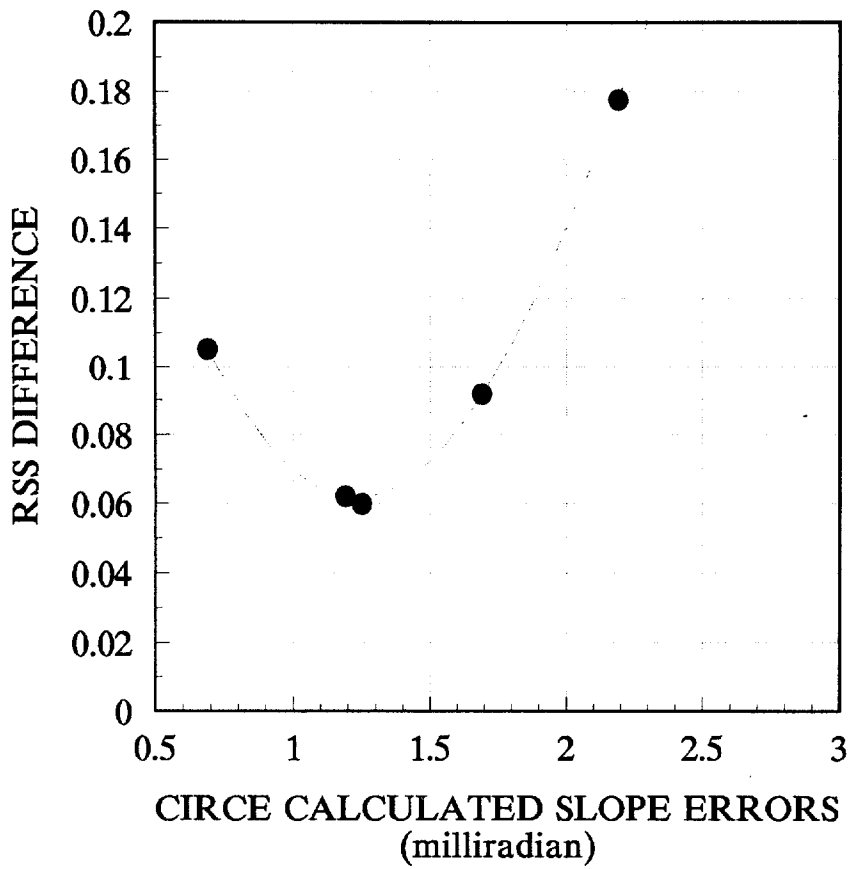
50 Suns/Contour



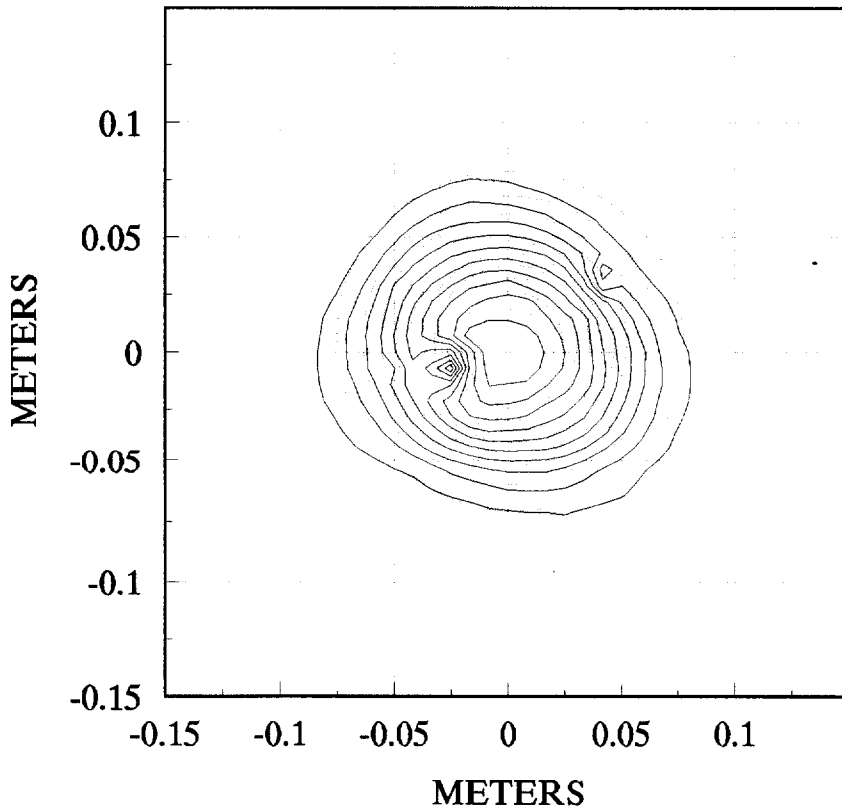
SKI 10.5 meters
 Peak Power 843 kW/sqm
 Peak Power Method Slope Error 1.5 mr
 RSS Differenc Method Slope Error 1.6 mr



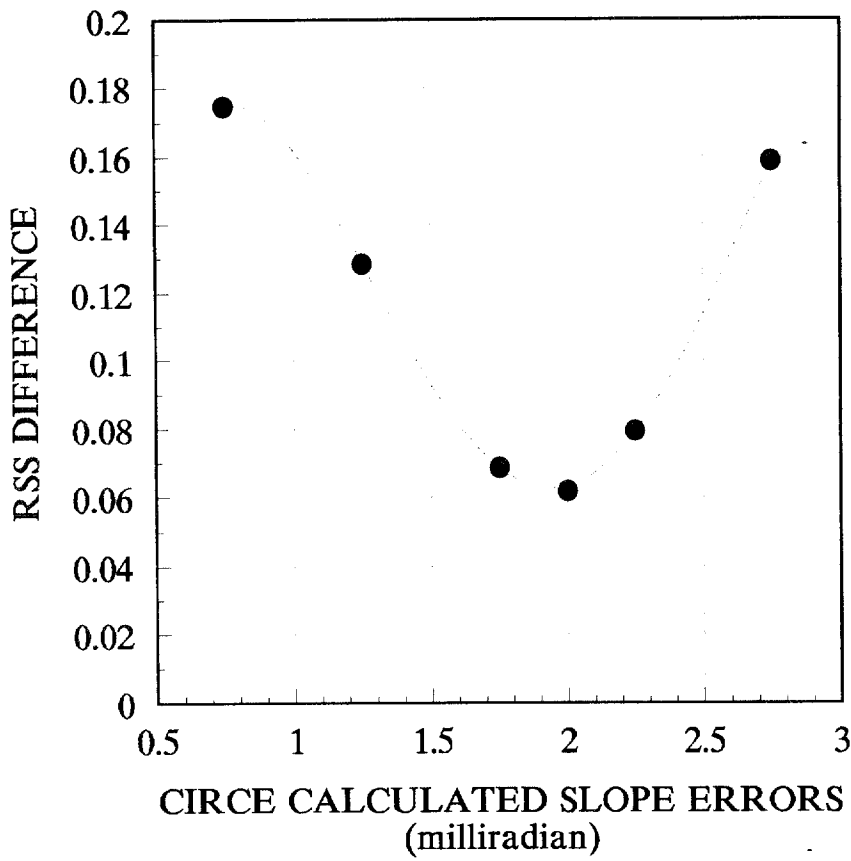
— Actual Image
 — CIRCE2 Prediction
 100 Suns/Contour



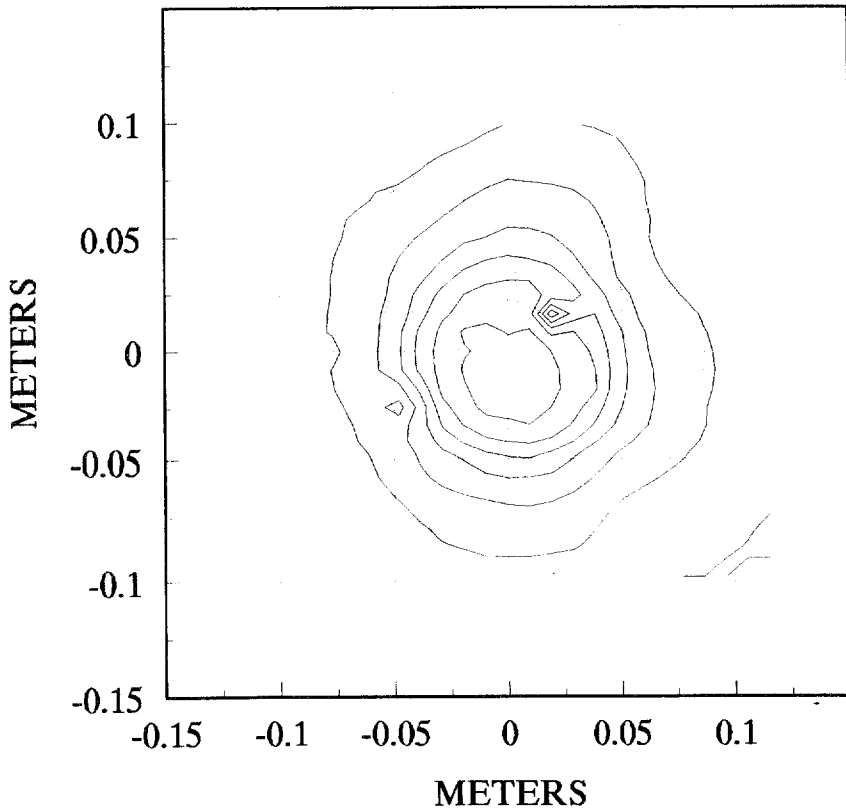
SKI 10 meters
 Peak Power 1186 kW/sqm
 Peak Power Method Slope Error 1.3 mr
 RSS Difference Method Slope Error 1.2 mr



— Actual Image
 — CIRCE2 Prediction
 100 Suns/Contour



SKI 9.5 meters
 Peak Power 755 kW/sqm
 Peak Power Method Slope Error 2.0 mr
 RSS Difference Method Slope Error 1.9 mr



— Actual Image
 — CIRCE2 Prediction
 100 Suns/contour

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