



# Comparison Between Results of the HELIOS and MIRVAL Computer Codes Applied to Central Receiver Solar-Energy Collection

C. N. Vittitoe, F. Biggs, P. L. Leary

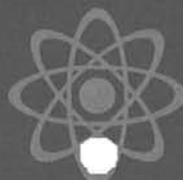
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COMPARISON BETWEEN RESULTS OF THE HELIOS AND MIRVAL COMPUTER CODES  
APPLIED TO CENTRAL RECEIVER SOLAR-ENERGY COLLECTION\*

Charles N. Vittitoe  
Frank Biggs  
Theoretical Division  
Sandia Laboratories, Albuquerque

Patricia L. Leary  
Applied Mathematics Division  
Sandia Laboratories, Livermore

ABSTRACT

The Sandia computer codes HELIOS and MIRVAL were developed to predict the optical performance of reflecting solar concentrators and to model power collection by central-receiver solar-energy power plants. HELIOS is an analytic code, whereas MIRVAL uses Monte Carlo ray-tracing techniques. They have been used both internally and externally in many studies including evaluation of heliostat-receiver design, parameter studies and safety analyses. The objective of this study was to verify that HELIOS and MIRVAL give the same performance predictions. The sample problem for comparison consists of a rectangular target and alt-azimuth heliostats deployed in a north field. The results indicate that HELIOS and MIRVAL closely agree on predictions of field performance and of power density on the target plane.

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## COMPARISON BETWEEN RESULTS OF THE HELIOS AND MIRVAL COMPUTER CODES APPLIED TO CENTRAL RECEIVER SOLAR-ENERGY COLLECTION

### Introduction

Harnessing solar power economically is a tremendous undertaking, with many possibilities for expensive waste. Computer codes have been developed to simulate various methods of collecting solar energy in order to avoid mistakes new technology can bring. The low cost of computer use compared with that of hardware development allows for many parametric studies prior to hardware decisions.

Because of the expense associated with high-accuracy hardware, solar-energy development seeks a trade-off between accuracy and cost. One of the largest expenses is in the heliostats (mirror array). In the optimum design, the cost (and therefore accuracy) is minimized subject to the constraint that we receive the desired power and the desired distribution of power on the receiver. Since computer models aid in making decisions involving large hardware expenditures, it is important that the models be verified comparatively and experimentally.

MIRVAL and HELIOS are two computer codes that are used to assess the overall optical performance of systems proposed for central receiver solar-energy collection. Documentation of the theory and use of these codes appears elsewhere.<sup>1-3</sup> Briefly, the inputs are the geometry of the heliostat field, tower, and receiver, along with miscellaneous physical data such as mirror reflectivity, insolation tables, etc., and the outputs are the thermal power through the receiver and the thermal power density on the plane of the receiver opening. Phenomena whose effects are simulated are shadowing, blocking, mirror tracking, random errors in tracking and in the conformation of the reflective surface, optical shape of the reflective surface, insolation, angular distribution of incoming sun rays to account for limb darkening and scattering, attenuation of light between the mirrors and the receiver, reflectivity of the mirror surface, and mirror-aiming strategy.

The methodologies of HELIOS and MIRVAL are quite different. MIRVAL is a Monte Carlo code. Rays of light are selected from the vicinity of the sun and are traced until they either enter the receiver or are lost in a prior absorption process or are deflected enough to miss the receiver. HELIOS, on the other hand, is an analytic code. The angular distribution of sunrays for the radiation incident on a concentrator is modified by convolution, using the fast Fourier transform, to incorporate the effects of other nondeterministic factors such as sun-tracking errors, surface slope errors, and reflectance properties.

MIRVAL has been used to compare sets of heliostat-field, tower-receiver, heliostat, and facet designs that have been proposed for the 10-MW electrical power tower now planned for Barstow, California. HELIOS has been used in the evaluation of individual heliostats at the Central Receiver Test Facility (CRTF) now in operation at Sandia Laboratories, Albuquerque, New Mexico. We tested several designs which had been proposed for the Barstow plant. MIRVAL and HELIOS have also been used in many other types of studies: safety analysis, parameter studies, power-tower-performance calculations, and comparison with experimental data.

Several previous checks of HELIOS were described in Chapter 9 of Reference 2. All of the quantitative comparisons with experiment were either for one heliostat or for one facet. Such comparisons give no validation to code features such as shadowing and blocking. Other work<sup>4</sup> examined consistency between prediction and experiment for a small field of 23 heliostats. As indicated elsewhere<sup>2</sup> detailed flux-density measurements are planned for the CRTF using large heliostat arrays and concurrent sunshape data. These data in turn will be used as experimental data for similar comparisons. This report examines the consistency of the two computer codes when applied to larger heliostat fields. We detail the basic input used to compare the codes, including the basic power-tower optical design. The code outputs described indicate that MIRVAL and HELIOS results are consistent.

Because often-used computer codes associated with a rapidly evolving technology are revised frequently to meet new demands, comparison is constrained to the versions of MIRVAL and HELIOS in use during July 1979. It is expected, however, that future versions of the codes will not alter the basic results or conclusions presented here.

### Code Input

As a result of interaction with a group from the Empresa Nacional de Ingenieria Y Tecnologia (INITEC) in Madrid, both MIRVAL and HELIOS were used to aid in the design of the 1-MW<sub>e</sub> CESA-1 (Central Energia Solar Almeria) solar-central receiver plant being built in Almeria, Spain. The preliminary design of the CESA-1 system (but with a spherical facet-surface shape) provided the details for the code comparison between HELIOS and MIRVAL.

The mirror array consists of 282 heliostats deployed in a north field, as shown in Figure 1. This figure was produced by HELIOS, as were all other computer graphics in this report. The coordinates of the heliostats are listed in the Appendix.

### HELIOSTAT LAYOUT IN EAST - NORTH PLANE

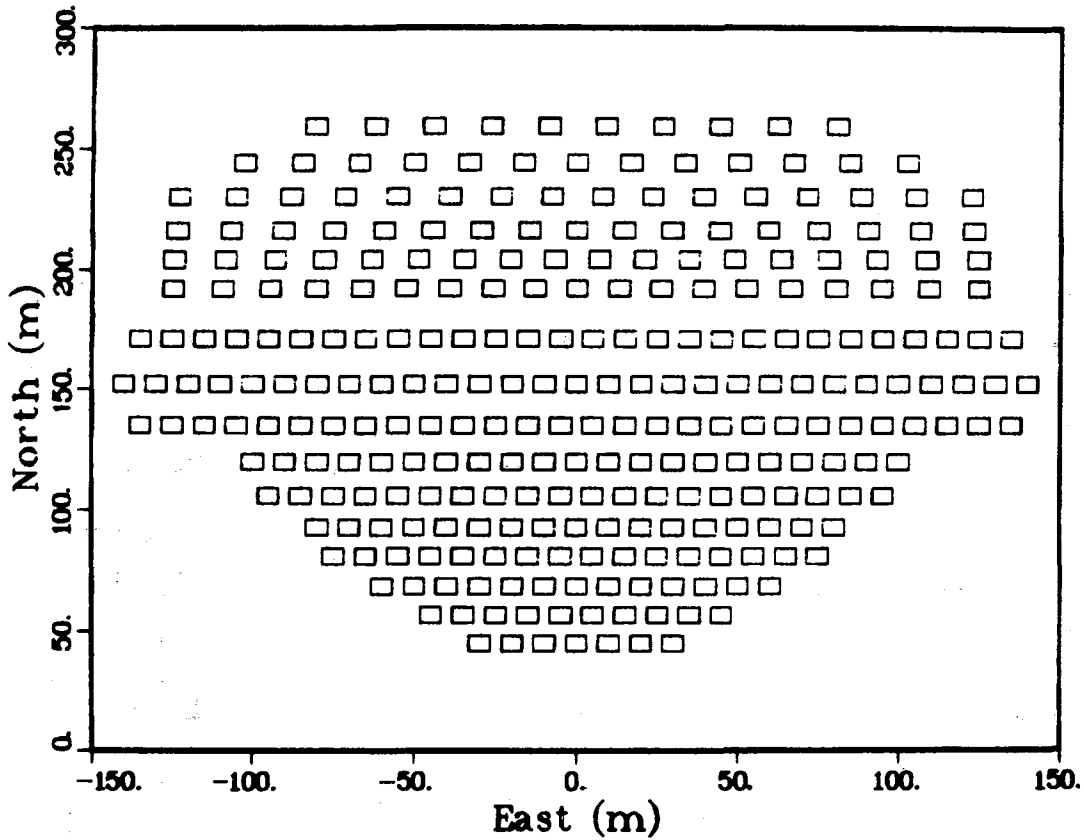


Figure 1. Preliminary Heliostat Layout in East-North Plane for the CESA-1 Solar Central-Receiver Project

All are aimed at the center of the target. Each heliostat consists of 10 panels arranged in a 2 x 5 pattern (Figure 2). Individual facets are canted to give optimum energy collection at noon on solar equinox. The facet surfaces are spherical with a reflectivity of 0.85. The focal length of each facet is determined from the projection P of the heliostat to tower-center distance onto a horizontal plane. The f values are listed in Table I.

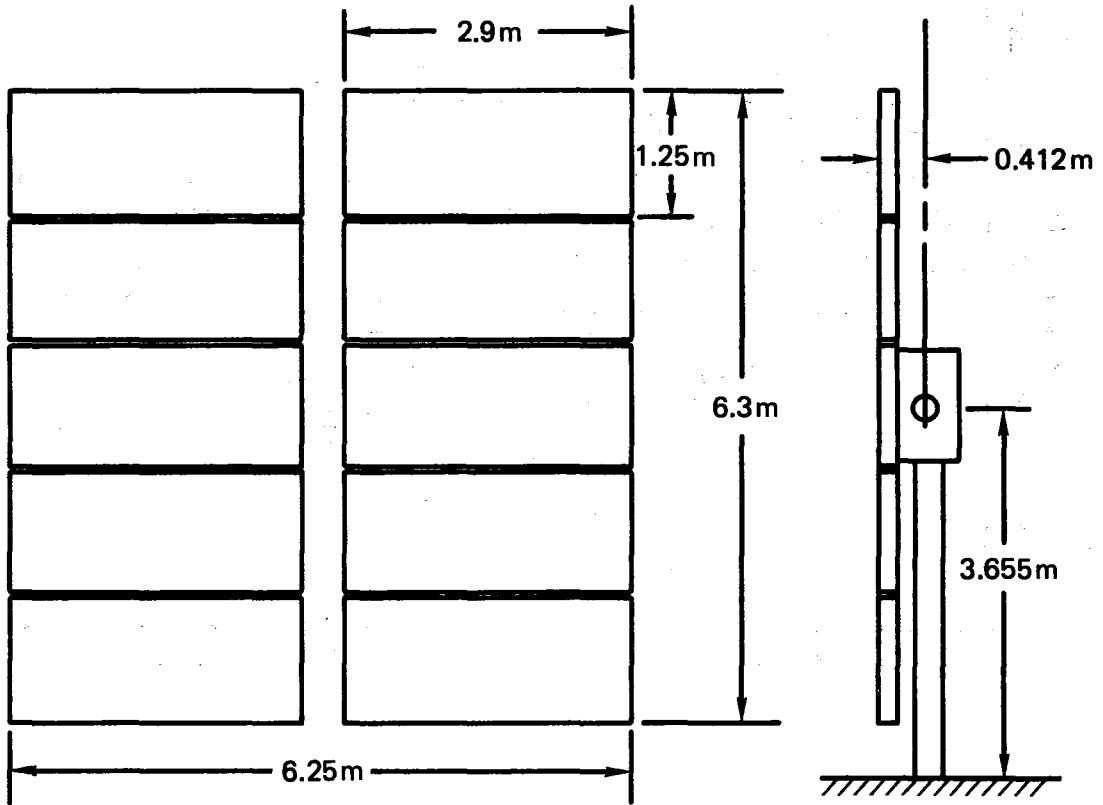


Figure 2. Heliostat Design



TABLE I  
FACET FOCAL LENGTHS

Projection P(m)	0-69	69-122	122-170	170-205	205-232	232-∞
f(m)	85	118	160	202	226	255

The target (Figure 3) is a 3.4 m square centered at 0 m east, 5 m north, and 60 m above the center of the tower base. The target is inclined 21.8° downward. The tower shadow is modeled as that cast by a cylinder 71 m tall with radius of 5.5 m.

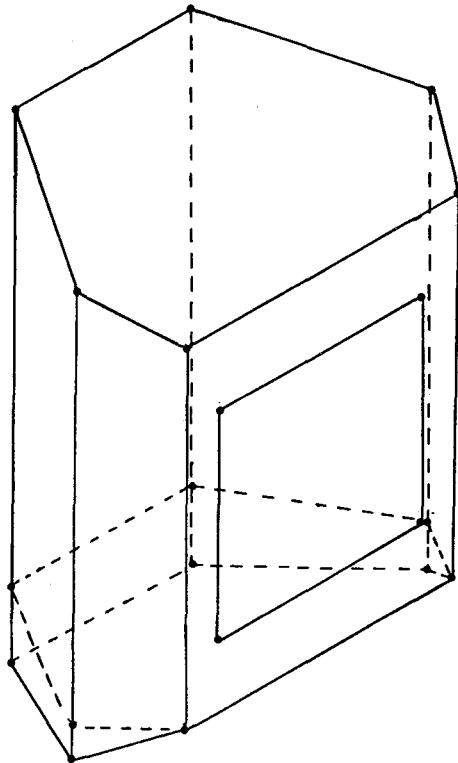


Figure 3. CESA-I Solar Receiver

The sun shape is taken as a uniform disk which subtends 9.29 mrad at the plant site. The calculation time is 10 A.M. on winter solstice. Insolation is  $700 \text{ W/m}^2$ . The latitude of the CESA-1 solar receiver site is  $37.099^\circ\text{N}$ . The atmospheric attenuation is modeled with Eq. (6.3-2) and (6.3-3) in Reference 2.

The dispersion in the error cone in HELIOS was set to 0.0033. In order to make the treatment of heliostat error sources consistent in the two codes, MIRVAL used 0.00165 rad for the standard deviation of the distribution function describing mirror-slope error and used no error in tracking angles. The users' guides<sup>1-3</sup> contain further description of the treatment of error sources.

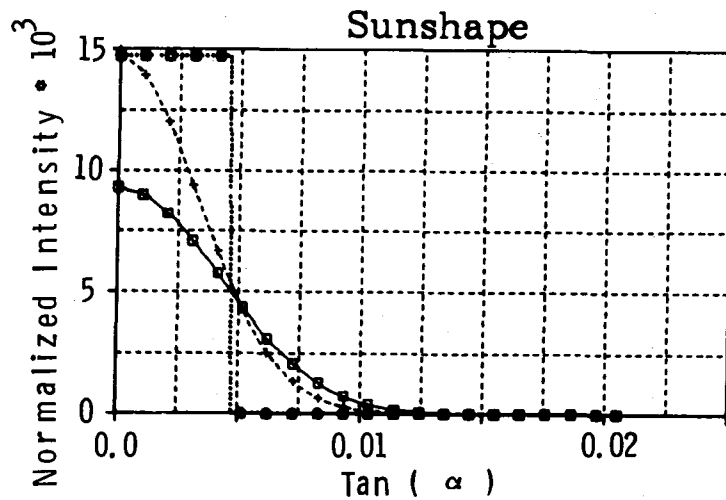


Figure 4. The uniform-disk sunshape (...), the error cone with a dispersion of 0.0033 (---), and the effective sunshape (—)

## Result Comparisons

The field performance parameters calculated by HELIOS and MIRVAL for the power-tower system defined in the preceding section are listed below.

### Power collected in aperture:

HELIOS 4.762 MW

MIRVAL 4.786 MW

### Error estimate

0.013 MW (rough estimate - see section 7.1 of reference 2)

0.018 MW (the probability is 0.97 that the estimate differs from the true value by less than 0.018 MW)

### Shadowing and blocking loss factor (area loss, not power loss at target):

HELIOS 0.92408

(.91888 for power loss at target)

MIRVAL 0.92541

(shadowing loss fraction 0.07149, blocking loss fraction 0.00334)

### Cosine loss factor (caused by angle of incidence at facets):

HELIOS 0.94910

MIRVAL 0.94913

### Spillage loss factor:

HELIOS 0.92174

MIRVAL 0.92142

### Atmospheric attenuation loss factor:

HELIOS 0.97388

MIRVAL 0.97329

### Flux density on target surface

This comparison is less straightforward since HELIOS calculates the flux density at a mesh of points while MIRVAL gives the average flux density in an

array of bins. Integration over portions of the target mesh in HELIOS gives the power distribution on the target within a set of bins. By using the area of the bins the average flux density within each bin is calculated. This can then be compared with the MIRVAL output as shown in Table II. The target coordinates in the table have their origin at the target center with the x axis eastward and the y axis tilted from the vertical direction 21.8° toward the north. Horizontal slices of data near the peak are compared in Figure 5. Comparison of vertical slices of data near the peak gives similar agreement. Consistency in the results remains when all heliostat error sources are set to zero. Figures 6 and 7 are graphs of the flux density calculated by HELIOS.

TABLE II  
 FLUX DENSITY DISTRIBUTION ON APERTURE (MW/m<sup>2</sup>)  
 FROM MIRVAL (HELIOS)

Target coordinates (in meters)						
x	1.70	1.02	0.34	-0.34	-1.02	-1.70
y						
1.70	.0721 (.0746)	.1685 (.1702)	.2163 (.2353)	.1668 (.1732)	.0717 (.0761)	
1.02	.1918 (.1953)	.5852 (.5612)	.9176 (.9165)	.5795 (.5770)	.1998 (.2016)	
0.34	.2841 (.2770)	.9666 (.9124)	1.6760 (1.6302)	.9633 (.9386)	.2992 (.2872)	
-0.34	.1826 (.1901)	.5474 (.5469)	.8650 (.8862)	.5612 (.5564)	.1960 (.1955)	
-1.02	.0664 (.0712)	.1531 (.1639)	.2118 (.2247)	.1582 (.1641)	.0694 (.0724)	
-1.70						

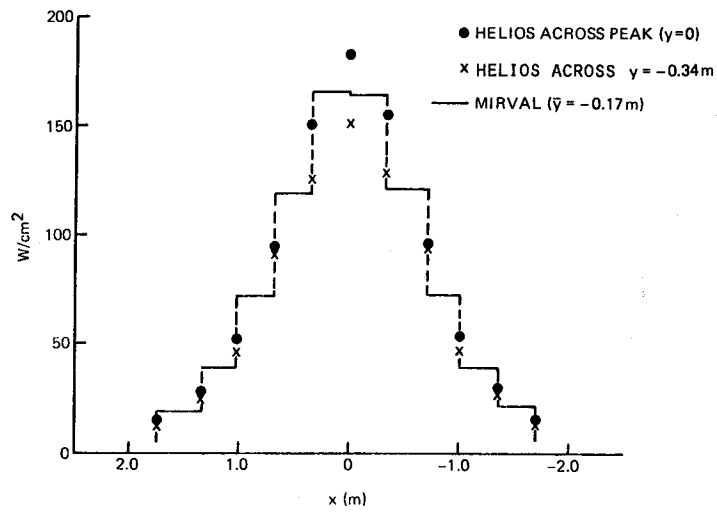


Figure 5. Horizontal Slices of Data Near Peak

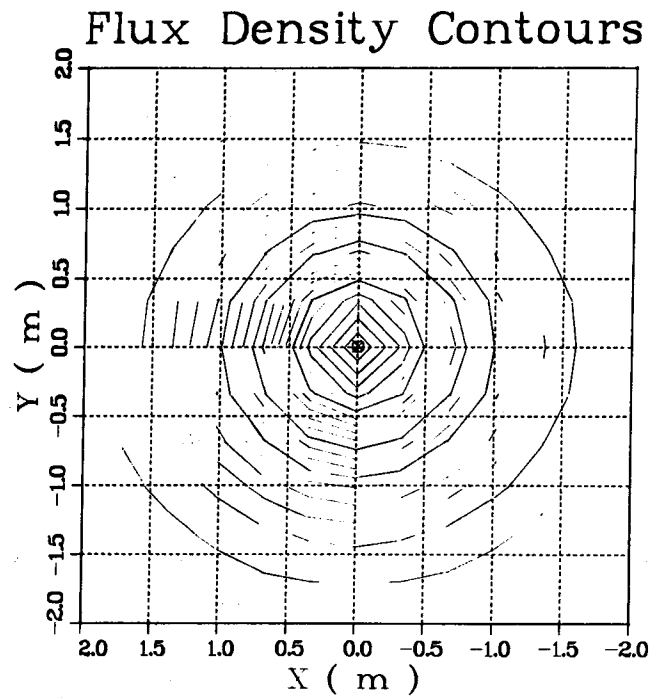
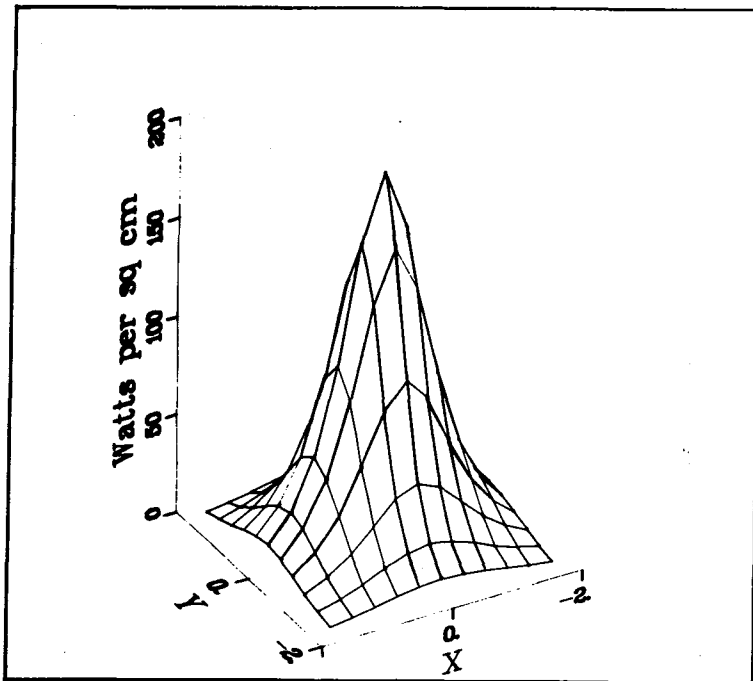


Figure 6. Isointensity contours of flux density on aperture at peak and  $(0.95 - 0.05I)$  for  $I = 0, 1, \dots, 18$ . Greater density of data points is needed for good contour resolution near the peak



Day = 355. Time = -2.0

Figure 7. Flux Density Distribution on Aperture

#### More Detailed Shadowing and Blocking Comparison

The extent of shadowing and blocking in this calculation is illustrated in Figures 8-9. The blocking diagram is a projection of the outer edges of the heliostats onto a unit sphere centered at the target center. Overlap indicates that the closer heliostat blocks a portion of the light from the overlapped heliostat. The shadowing diagram is a projection of the outer edges of the heliostats and a projection of the receiver-tower model onto a plane orthogonal to the sun's rays. The cross-hatched section represents the tower. Overlap here indicates that the closer heliostat or the tower shadows a portion of a heliostat farther from the tower. Figures 8 and 9 show little blocking or shadowing other than the tower shadowing.

Early calculations indicated that the blocking results of MIRVAL and HELIOS were not consistent. A time shift to 4 P.M. (to emphasize the effect) and concentration upon heliostat number 136 (see Appendix) indicated this heliostat was being blocked by heliostat number 107. MIRVAL predicted the blocked fraction of the heliostat reflective area to be 0.0165. The corresponding HELIOS prediction was 0.042. Shortly before this discovery workers at INITEC<sup>5</sup> indicated an inconsistency between a HELIOS prediction for blocking and a hand calculation for this heliostat-field layout. Review of

the code revealed the intent to locate the center of the unit sphere used for blocking at the center of the target. A programming error prevented these coordinates from being set, causing use of default CRTF parameters and thus the inconsistency. Code correction on July 5, 1979, made the MIRVAL, HELIOS and INITEC blocking results consistent. The shadowing and blocking diagrams for heliostats influencing number 136 at 4 P.M. are indicated in Figure 10.

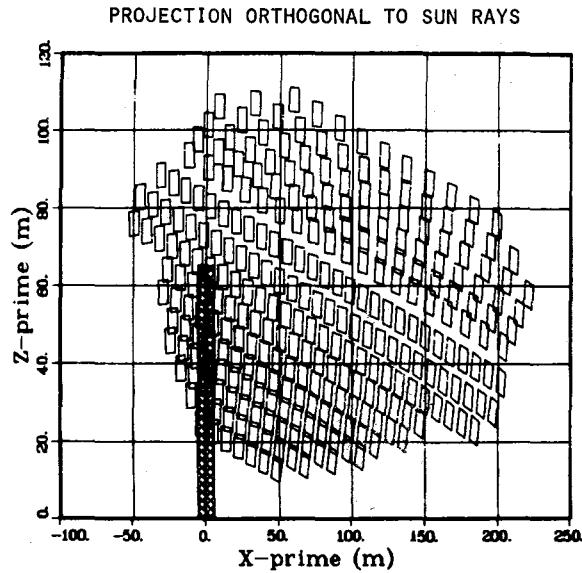


Figure 8. The Blocking Diagram

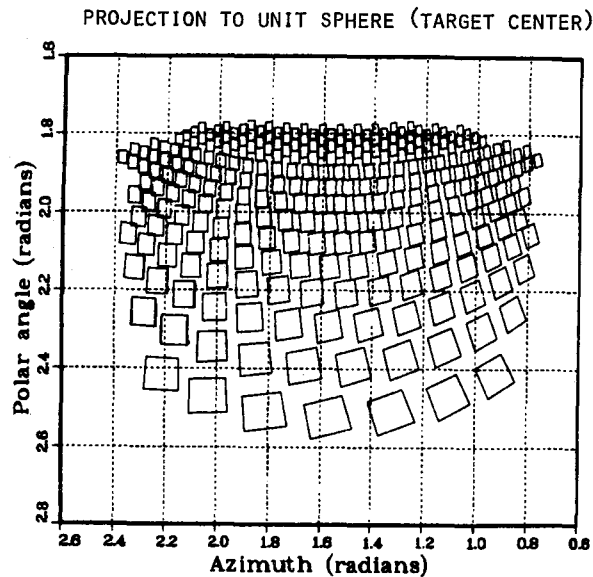


Figure 9. The Shadowing Diagram

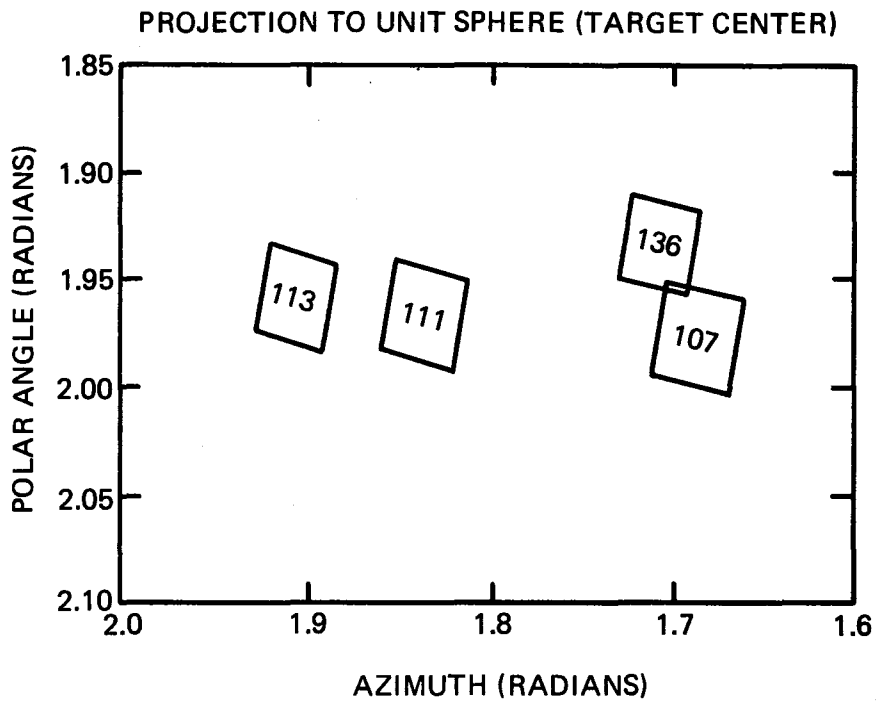
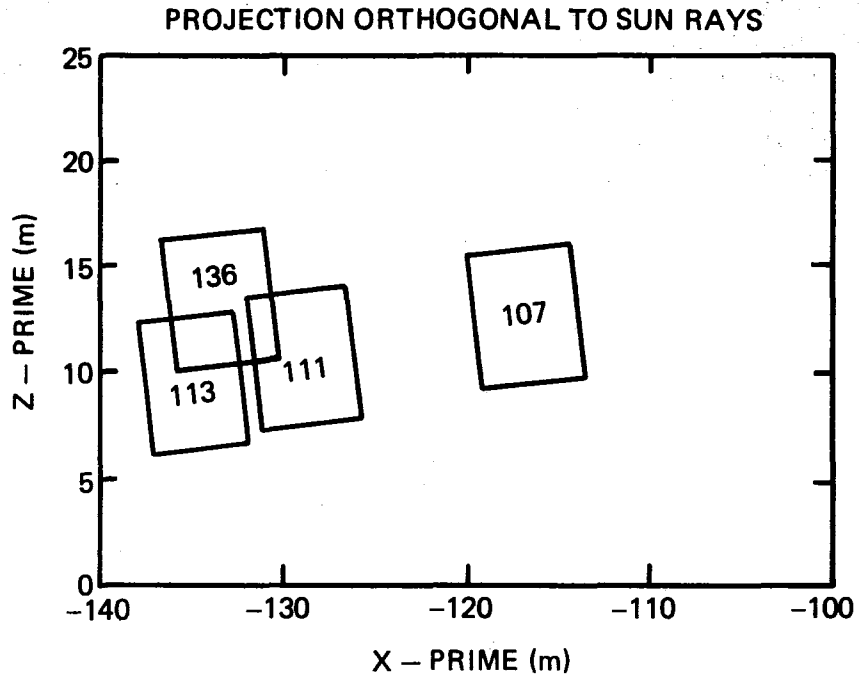


Figure 10. Shadowing and Blocking Diagrams Examining Heliostat Number 136 at 4 P.M. on Winter Solstice - The Top Graph Indicates Shadowing; The Bottom Blocking



## Conclusions

With a uniform disk (pillbox) sunshape, and the preliminary power-tower design for the CESA-1 central receiver, the HELIOS-MIRVAL results are in good agreement. Descriptions of heliostat error sources can be selected in a consistent manner. No differences are apparent in predictions of the two computer codes. Since the codes are based upon widely different approaches (cone-optics vs Monte Carlo) this consistency indicates a measure of validity in the two approaches. Comparison with experimental data for a large field remains. Nevertheless, we have greater confidence in the codes because of this exercise.



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